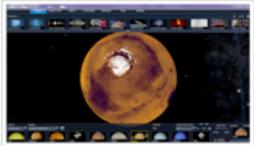
WorldWide Telescope Ambassadors Program http://www.cfa.harvard.edu/WWTAmbassadors/

Harvard University, WGBH & Microsoft Research Alyssa Goodman, Patricia Udomprasert, Annie Valva & Curtis Wong



What is WorldWide Telescope and its Ambassadors Program? WorldWide Telescope (WWT) is a fantastic "Universe Information System"

created primarily by Curtis Wong and Jonathan Ray at Microsoft Research. It functions as a Virtual Astronomical Observatory linking its users to much of the world's store of online data and information about our Universe. WWT is evolving to become a key research tool within the online astronomy ecosystem known in the US presently as the "VAO" (see A. Goodman's "Seamless Astronomy" talk at this meeting), but it also offers unprecedented new opportunities for STEM outmach.

The WorldWide Telescope Ambassadors Program promotes WWT as a future-learning way to teach and learn STEM concepts by recruiting astronomically-literate volunteers who are trained to be experts in using WWT as a teaching tool.

Who are we?

Our current collaboration brings together professional astronomers and science educators at Harvard, computational virtueses at MS Research, and STEM education and outreach specialists at WGBH. The next phase of the project (see table below) will include participants from selected areas within the US, including Washington, Florida, Arizona, Alaska, and Appalachia.

Who are the WWT Ambassadors, and what do they do?

WWT Ambassadors are carefully recruited for training from amongst: 1) retired STEM professionals and amateur astronomers with a demonstrable deep knowledge of astronomy and physics; 2) undergraduate and graduate students and postdoctoral fellows in Astronomy and Physics; and 3) science teachers. In their training, Ambassadors learn how to use WWT's tools in general, and also how to create and publish guided "Tours" of astrophysical concepts. These Tours allow users to display beautiful astronomical images in their proper context in the night sky, while demonstrating the physical principles at work in those images. Ambassadors can create and use materials within WWT; give volunteer presentations at variety of public venues; help out in classroom settings; or choose to do more than one of the above!



What's the whole plan, and what are the program's goals?

We are presently preparing a proposal to the National Science Foundation, based in large part on our "Pilot" experience, to implement "Phase I" of the Ambassadors Project (see table), where we will begin a limited expansion within the US, carefully selecting cities and partners where we will be able to maximize success with the available resources, while increasing the socioeconomic diversity of our sites. We plan to expand nationally in Phase II, and internationally in Phase III. With minimal advertising, we have already received inquiries from dozens of interested and qualified potential volunteers in multiple states and countries.

A critical goal of this project is to create a **full astronomy curriculum using WWT Tours created by our Ambassadors**. These Tours will be vetted by the astronomy and science education professionals within our collaboration, and they will be freely available, centrally managed, and searchable, through web services at WGBH. The entire WWT Ambassadors "Tour Curriculum" will be integrated with WGBH Teachers' Domain, which currently has nearly 400,000 registered users.

WorldWide Telescope can help change how students learn science by demonstrating the joys of inquiry and discovery, and the WWT Ambassadors Program is designed to help to increase science literacy in the general public while forming intergenerational connections within their communities.



E/PO ADVERTISEMENT

The WorldWide Telescope Ambassadors Program

www.cfa.harvard.edu/WWTAmbassadors/

A A + C http://www.cfa.harvard.edu/WWTAmbassadors/

WorldWide Telescope Ambassadors Program

About Galileo Tour Project Team How to get involved Tour-making Tutorials Documents Events Protected WorldWide Telescope







See a video of our interactive Tour in WWT ecreating Galileo's historic observations of Jupiter's moons.

WorldWide Telescope (WWT) is a rich visualization environment that functions as a virtual telescope, allowing anyone to make use of professional astronomical data to explore and understand the universe. As of early 2010, the new WWT Ambassadors Program is recruiting astronomicallyliterate volunteers, including retired scientists engineers-all of whom will be trained to be experts in using WWT as a teaching tool. Ambassadors will give volunteer presentations at public libraries, community centers, museums, and schools, demonstrating WWT's power to help laypeople visualize and understand our universe. Ambassadors will learn how to create and publish guided "tours" of astrophysical concepts, which allow users to display beautiful astronomical images in their proper context in the night sky, while demonstrating the physical principles at work in those images.

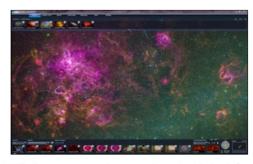
Tour creators will be able to draw upon and link tours to highly vetted multimedia content from NOVA, the renowned PBS multi-platform series produced by WGBH. Virtual tours will be freely available and centrally managed in order to form a comprehensive astronomy curriculum for both formal and informal educational use. The tours will be searchable and distributed online from popular websites such as NOVA Online and WGBH Teachers' Domain, touting almost 400,000 registered users. [www.teachersdomain.org]

WWT Ambassadors will help to increase science literacy in the general public while forming intergenerational connections within their communities.

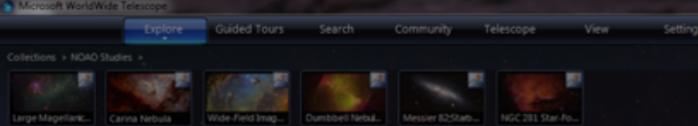


C Q- Google

6th grade students at Clarke Middle School, Lexington, MA learn about the universe using the WorldWide Telescope



WWT allows users to explore our universe in rich detail, from our solar system out to the largest observed structures in the cosmos.

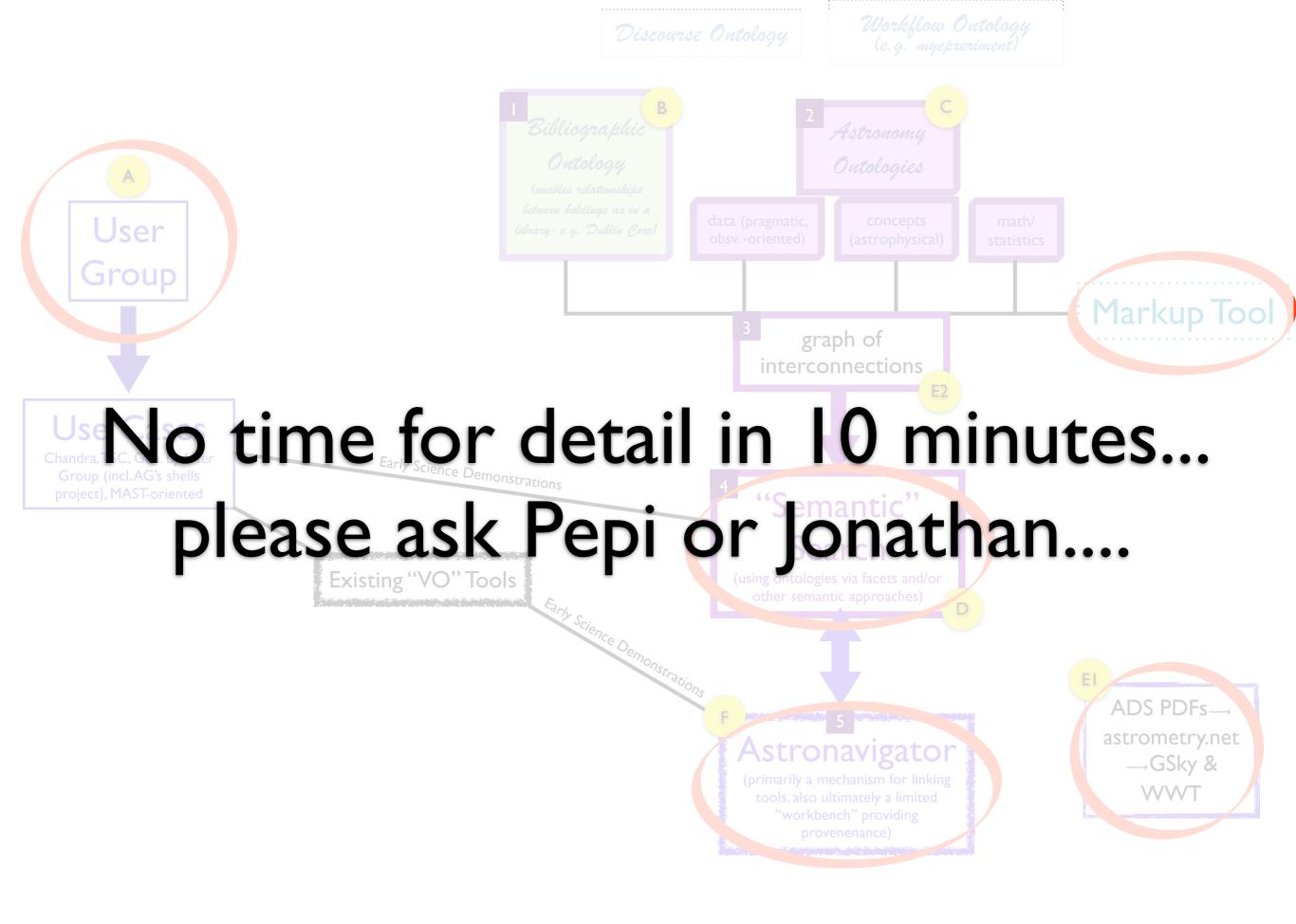


Seamless Astronomy

Alyssa A. Goodman Harvard-Smithsonian Center for Astrophysics

with Alberto **Accomazzi**, Rahul **Davé**, Gus **Muench** & Michael **Kurtz** (Harvard-Smithsonian CfA); Tim **Clark** (Massachusetts General Hospital/Harvard Medical School); Jonathan **Fay** & Curtis **Wong** (Microsoft Research)

+extended & upcoming collaboration with Chris Borgman & Alberto Pepe* (UCLA);
 Doug Burke; Sarah Block, Pepi Fabbiano, et al. (CfA); E. Bressert (U. Exeter);
 J. Hendler & D. McGuinness (RPI); A. Conti & C. Christian (STScI); A. Connolly et al. (U. Washington)



Realm of Seamless Astronomy





What should we do now?

Focus on interconnections/modularity

interconnections/modularity

Involve more "free agents" free agents

Provide institutional data/code repositories

repositories

Make users aware w/o "software-speak"

awareness/usability

This simple argument, first made at the 2009 WWT session at AAS, seems to be working:

"Astronomy research tools should work as seamlessly as travel research tools."

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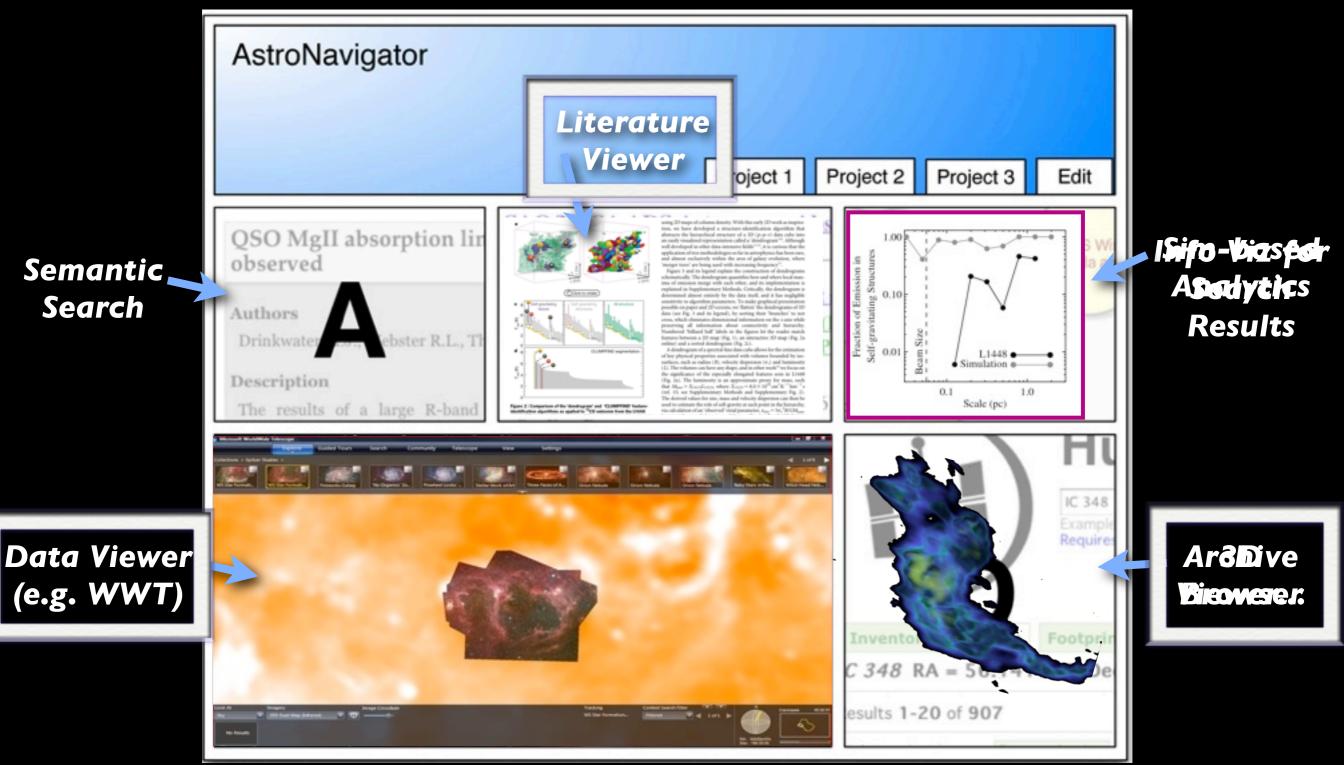
Distorted galaxy NGC 2442 can be found in the southern constellation of the flying fish, (Piscis) Volans. Read More

Welcome, Alyssa Goodman logout What's on your mind? Share Share Share Image: Starred Image: Starred Image: Starred Image: Starred	Facebook	
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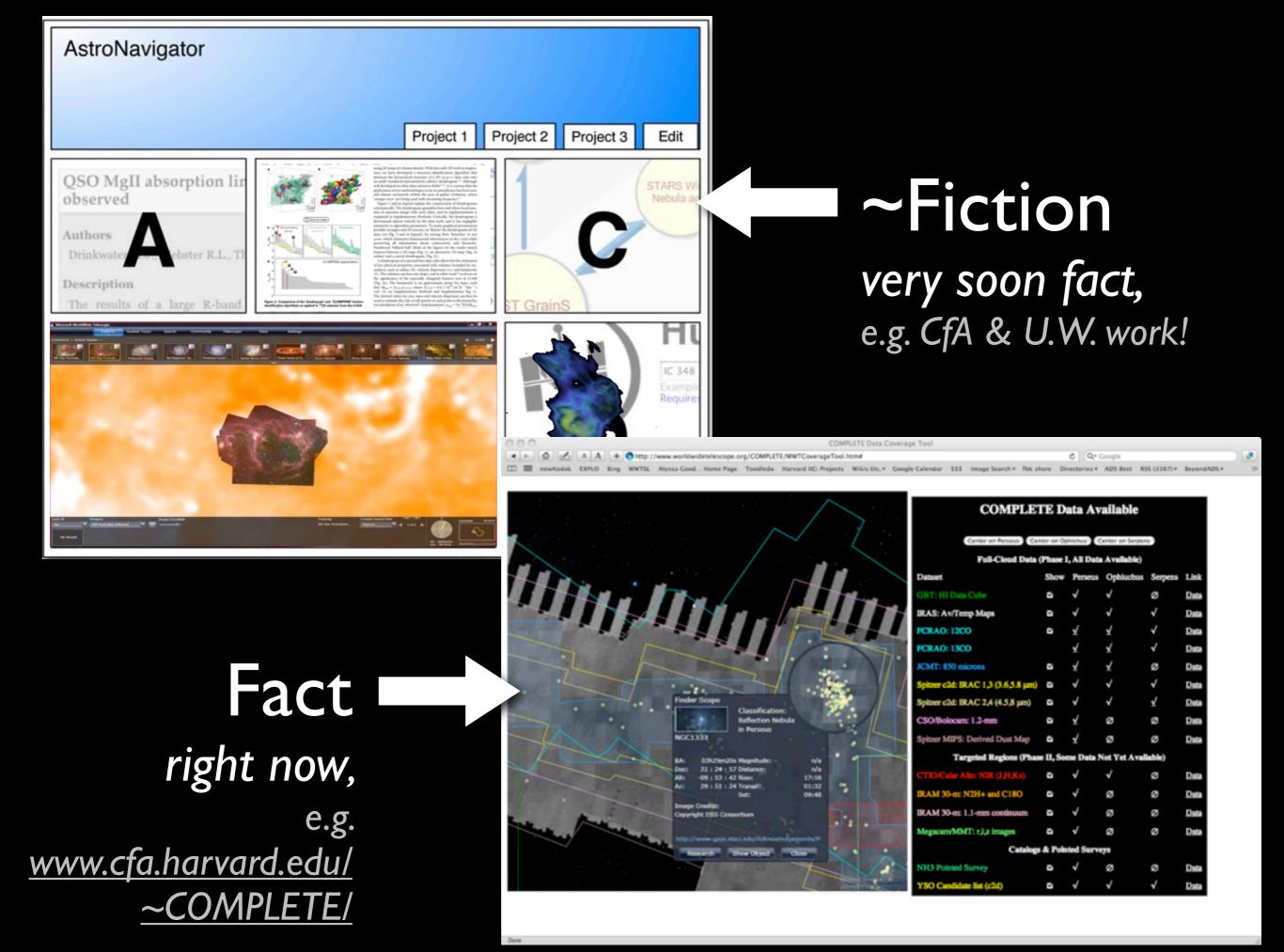
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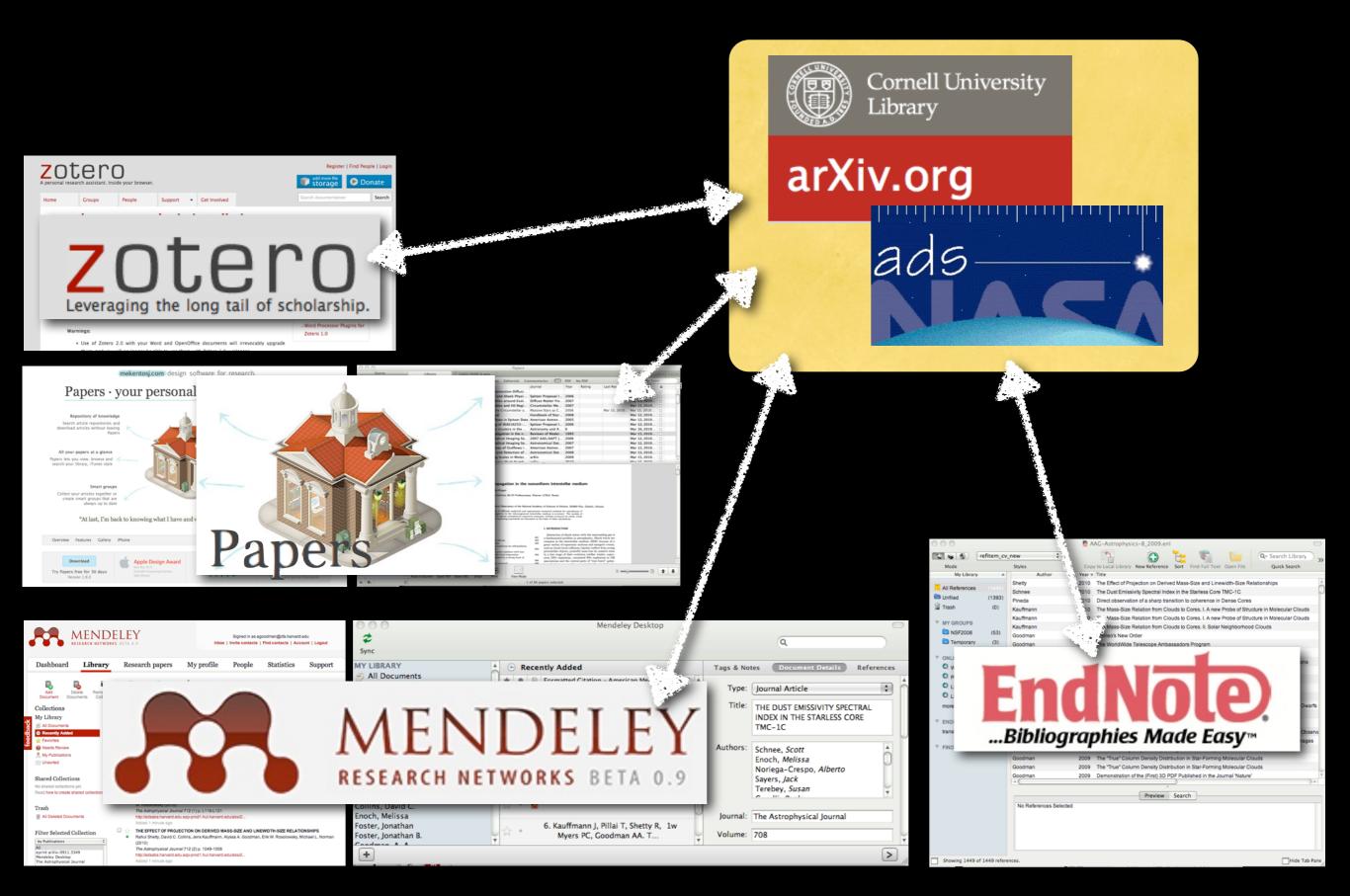
Seamless Astronomy



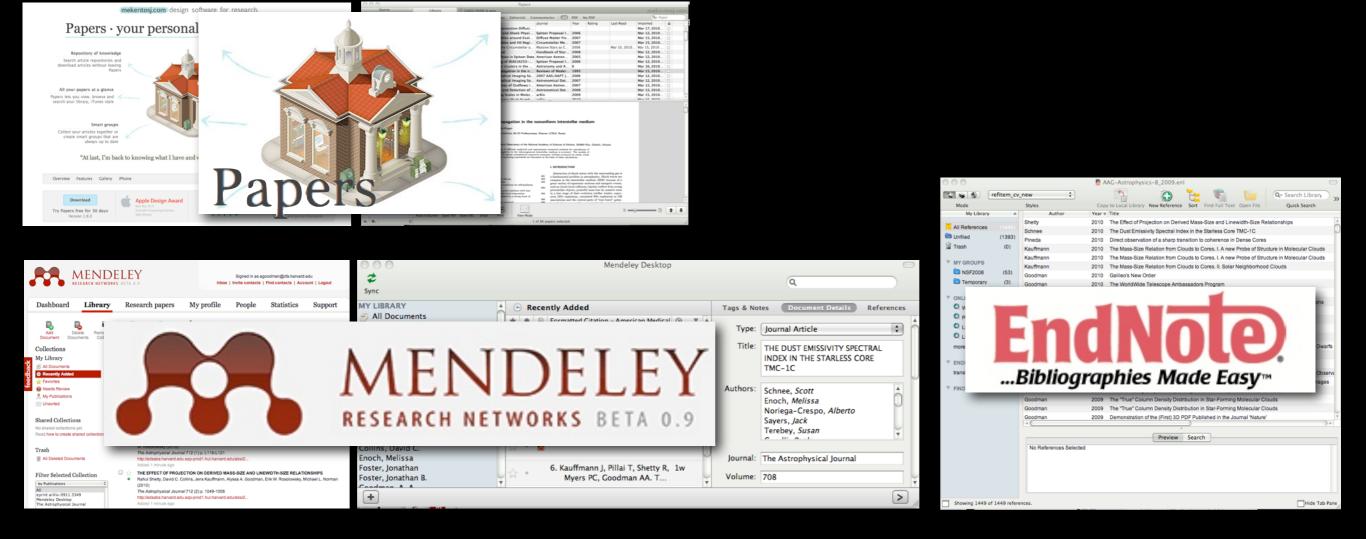
Mockup based on work of Eli Bressert, excerpted from NASA AISRP proposal by Goodman, Muench, Christian, Conti, Kurtz, Burke, Accomazzi, McGuinness, Hendler & Wong, 2008



Literature Handling: Diverse Apps, Common Data



What fraction of astronomy researchers know about these tools?



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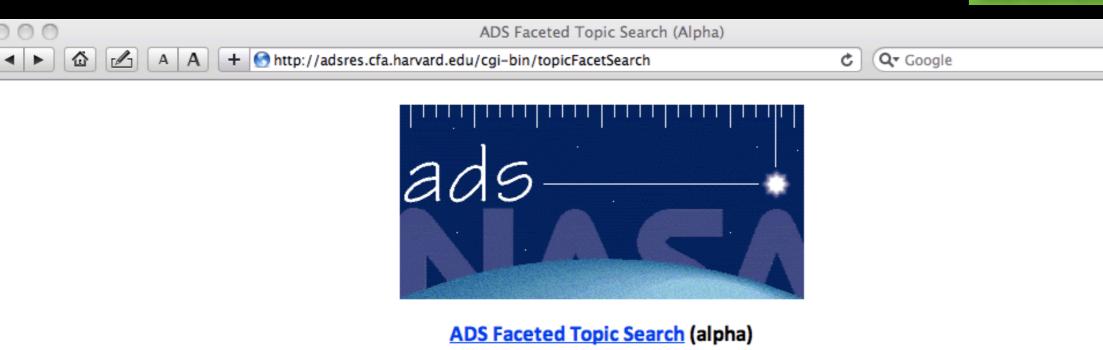
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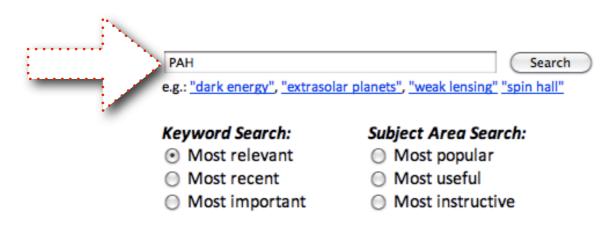
Leveraging the long tail of scholarship.

= Use of Zotero 2.0 with your Word and OpenOffice d

interconnections/modularity







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"alpha" Faceted Topic Search in ADS (courtesy of Michael Kurtz & Alberto Accomazzi) ٠.

ADS Query Results

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SAO/NASA Astrophysics Data System (ADS)

Query Results from the ADS Database

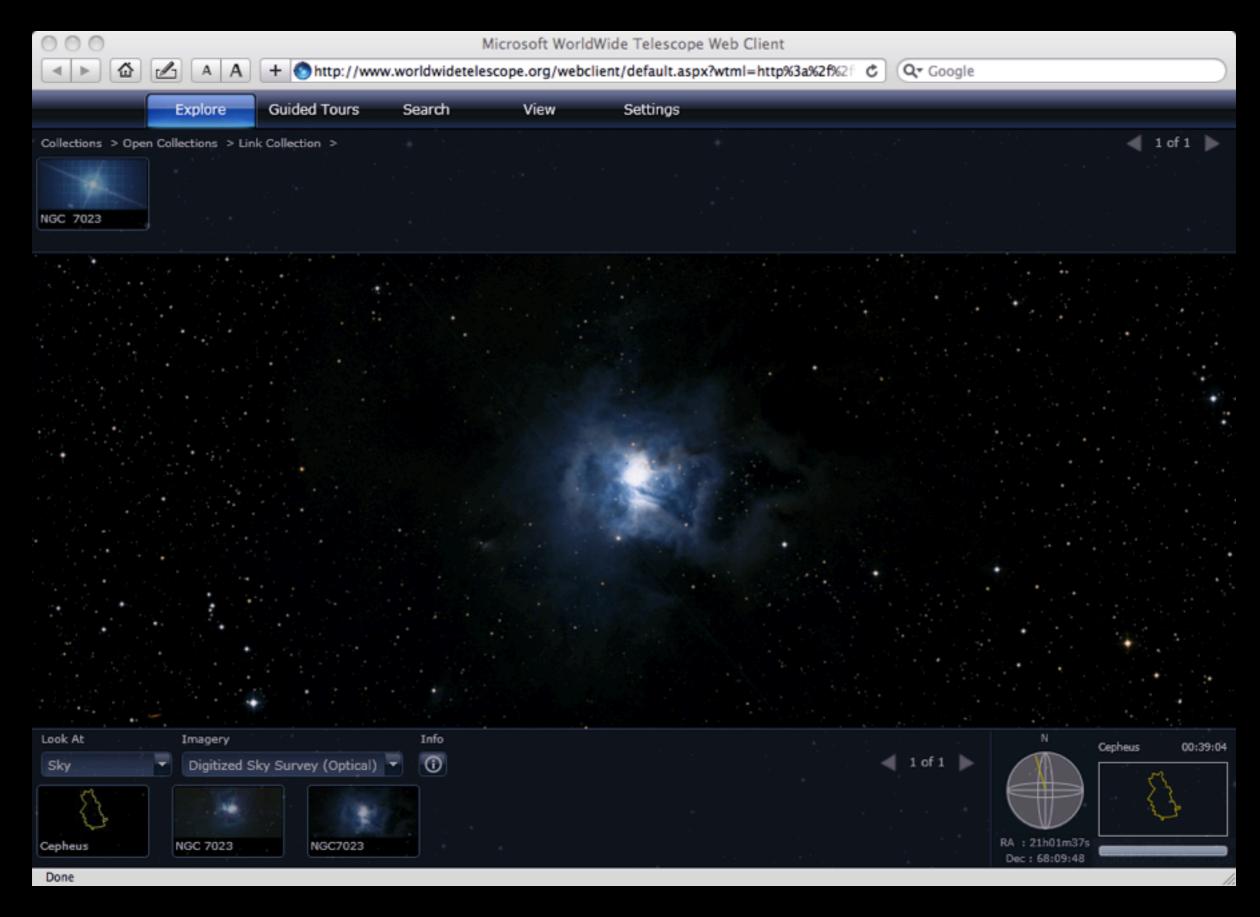
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<u>SBSG 0335-052 (4)</u> <u>QSO B2300+086 (4)</u> <u>NGC 7331 (4)</u> <u>NGC 4151 (4)</u> <u>NGC 1808 (4)</u> <u>NGC 1097 (4)</u> <u>NAME CAMPBELL'S HYDROGEN</u> STAR (4)	3 Description 2007ApJ654L495 Spoon, H. W. W.; Marshall, J. A.; Houck, J. R.; Elitzur, M.; Hao, L.; Armus, L.; Brandl, B. R.; Charmandaris, V.	95.232 Jan 2007 <u>A E E X R C</u> Mid-Infrared Galaxy Classification Based on Silicate Obso Width	
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list of objects with links to WWT browser (thanks to ADS team & Jonathan Fay)

Go to bottom of page

4

Now we got to NGC 7023 by using the literature as a filter.



interconnections/modularity

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Image ssc2003-06f

http://www.spitzer.caltech.edu/Media/releases/ssc2003-06/ssc200 ~ 🗘

usually buried within dark clouds. They are formed when supersonic gas ejected from a forming protostar, or embryonic star,

The Spitzer image was obtained with the infrared array camera. Emission at 3.6 microns is shown as blue, emission from 4.5

ACIAT is a attilize anomal of a low mass protector significant and areating a bipalar artur aided autilow. The control

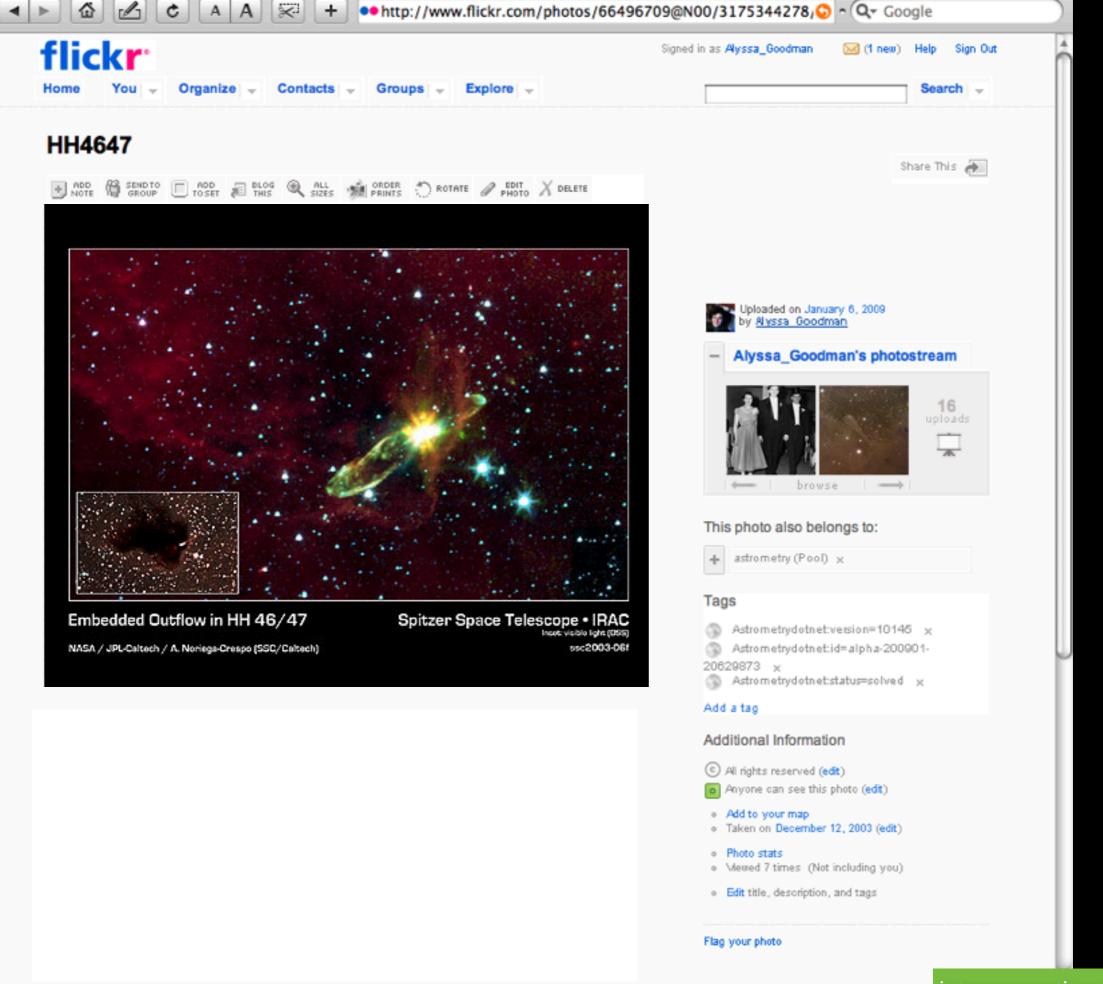
interacts with the surrounding interstellar medium. These young stars are often detected only in the infrared.

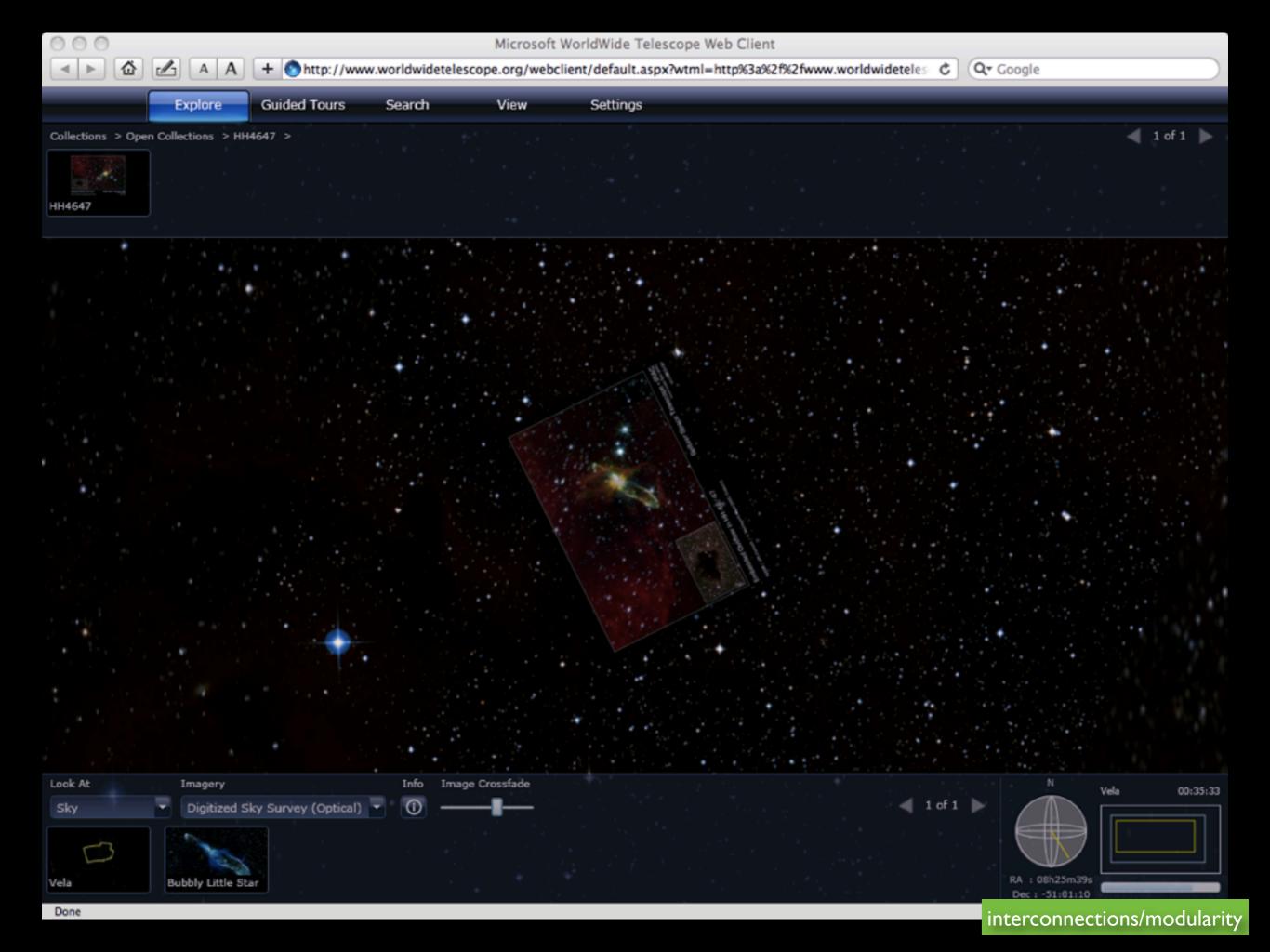
and 5.8 microns has been combined as green, and 8.0 micron emission is depicted as red.

Spitzer Space Telescope NASA Newsroom Podcasts Features About Spitzer Home Images NEWSROOM **Press Releases** INTRODUCTION PRESS RELEASE VISUALS QUICK FACTS Chronological By Subject - Outside Institutions What's Happening Archive Visuals - Image Use Policy Update Notifications Mailing List RSS Feed (XML) References Fast Facts - Press Kit (.pdf) Fact Sheet (.pdf) Field Guides Glossary Media Contacts Embedded Outflow in HH 46/47 Spitzer Space Telescope • IRAC NASA / JPL-Caltech / A. Noriega-Crespo (SSC/Caltech) 30.000-Credit: NASA/JPL-Caltech/A. Noriega-Crespo (SSC/Caltech), Digital Sky Survey HH46/47 This image from NASA's Spitzer Space Telescope transforms a dark cloud into a silky translucent veil, revealing the molecular outflow from an otherwise hidden newborn star. Using near-infrared light, Spitzer pierces through the dark cloud to detect the embedded outflow in an object called HH 46/47. Herbig-Haro (HH) objects are bright, nebulous regions of gas and dust that are

Seamlessness through...

flickr + astrometry.net + WWT !?





Coming (Very) Soon...

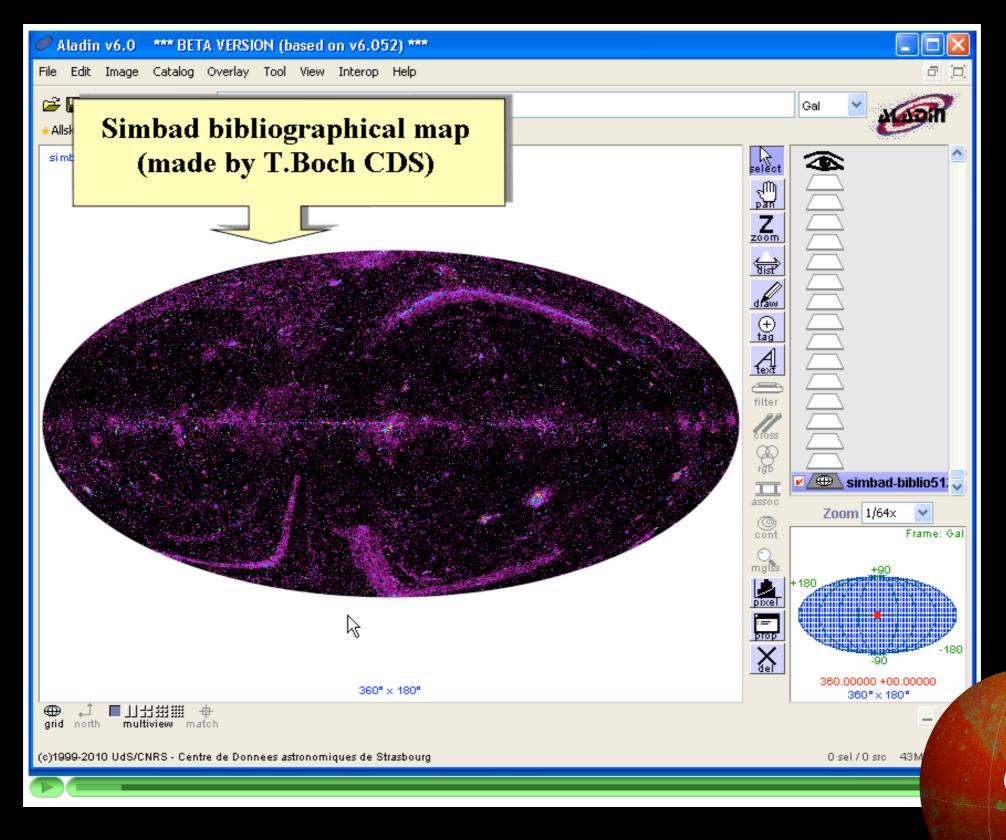
Historical Image Layer Extracted from ALL ADS holdings (using astrometry.net)

[ADS+WWT+who wants to help?]

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[ADS+CDS+WWT are doing it!]

Prototype of Articles on the Sky (April 2010)



with thanks to CDS/Pierre Fernique

The future is here... data IN articles

interconnections/modularity

free agents

repositories

awareness/usability

Note: This work came from the "'AstroMed" project am.iic.harvard.edu



LETTERS

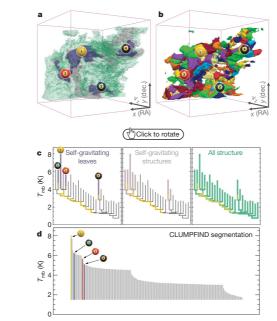


Figure 2 Comparison of the 'dendrogram' and 'CLUMPFIND' feature identification algorithms as applied to ¹³CO emission from the L1448 region of Perseus. a, 3D visualization of the surfaces indicated by colours in the dendrogram shown in c. Purple illustrates the smallest scale selfgravitating structures in the region corresponding to the leaves of the dendrogram; pink shows the smallest surfaces that contain distinct selfgravitating leaves within them; and green corresponds to the surface in the data cube containing all the significant emission. Dendrogram branches corresponding to self-gravitating objects have been highlighted in yellow over the range of T_{mb} (main-beam temperature) test-level values for which the virial parameter is less than 2. The x-y locations of the four 'selfgravitating' leaves labelled with billiard balls are the same as those shown in Fig. 1. The 3D visualizations show position-position-velocity (p-p-v) space. RA, right ascension; dec., declination. For comparison with the ability of dendrograms (c) to track hierarchical structure, d shows a pseudodendrogram of the CLUMPFIND segmentation (b), with the same four labels used in Fig. 1 and in a. As 'clumps' are not allowed to belong to larger structures, each pseudo-branch in d is simply a series of lines connecting the maximum emission value in each clump to the threshold value. A very large number of clumps appears in **b** because of the sensitivity of CLUMPFIND to noise and small-scale structure in the data. In the online PDF version, the 3D cubes (**a** and **b**) can be rotated to any orientation, and surfaces can be turned on and off (interaction requires Adobe Acrobat version 7.0.8 or higher). In the printed version, the front face of each 3D cube (the 'home' view in the nteractive online version) corresponds exactly to the patch of sky shown in Fig. 1, and velocity with respect to the Local Standard of Rest increases from front (-0.5 km s^{-1}) to back (8 km s^{-1})

data, CLUMPFIND typically finds features on a limited range of scales, above but close to the physical resolution of the data, and its results can be overly dependent on input parameters. By tuning CLUMPFIND's two free parameters, the same molecular-line data set⁸ can be used to show either that the frequency distribution of clump mass is the same as the initial mass function of stars or that it follows the much shallower mass function associated with large-scale molecular clouds (Supplementary Fig. 1).

Four years before the advent of CLUMPFIND, 'structure trees'⁹ were proposed as a way to characterize clouds' hierarchical structure

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NATURE Vol 457 1 January 2009

using 2D maps of column density. With this early 2D work as inspiration, we have developed a structure-identification algorithm that abstracts the hierarchical structure of a 3D (p–p–v) data cube into an easily visualized representation called a 'dendrogram'¹⁰. Although well developed in other data-intensive fields^{11,12}, it is curious that the application of tree methodologies so far in astrophysics has been rare, and almost exclusively within the area of galaxy evolution, where 'merger trees' are being used with increasing frequency¹³.

Figure 3 and its legend explain the construction of dendrograms schematically. The dendrogram quantifies how and where local maxima of emission merge with each other, and its implementation is determined almost entirely by the data itself, and it has negligible sensitivity to algorithm parameters. To make graphical presentation possible on paper and 2D screens, we 'flatten' the dendrograms of 3D data (see Fig. 3 and its legend), by sorting their 'branches' to not cross, which eliminates dimensional information on the x axis while preserving all information about connectivity and hierarchy. Numbered 'billiard ball' labels in the figures 1b the reader match features between a 2D map (Fig. 1), an interactive 3D map (Fig. 2a online) and a sorted dendrogram (Fig. 2c).

A dendrogram of a spectral-line data cube allows for the estimation of key physical properties associated with volumes bounded by isosurfaces, such as radius (R), velocity dispersion (σ_v) and luminosity (L). The volumes can have any shape, and in other work¹⁴ we focus on the significance of the especially elongated features seen in L1448 (Fig. 2a). The luminosity is an approximate proxy for mass, such that $M_{\text{lum}} = X_{13\text{CO}}L_{13\text{CO}}$, where $X_{13\text{CO}} = 8.0 \times 10^{20} \text{ cm}^2 \text{ K}^{-1} \text{ km}^{-1} \text{ s}$ (ref. 15; see Supplementary Methods and Supplementary Fig. 2). The derived values for size, mass and velocity dispersion can then be used to estimate the role of self-gravity at each point in the hierarchy, via calculation of an 'observed' virial parameter, $\alpha_{obs} = 5\sigma_v^2 R/GM_{hum}$ In principle, extended portions of the tree (Fig. 2, yellow highlighting) where $\alpha_{obs} < 2$ (where gravitational energy is comparable to or larger than kinetic energy) correspond to regions of p-p-v space where selfgravity is significant. As α_{obs} only represents the ratio of kinetic energy to gravitational energy at one point in time, and does not explicitly capture external over-pressure and/or magnetic fields16, its measured value should only be used as a guide to the longevity (boundedness) of any particular feature.

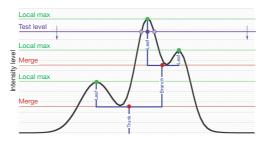


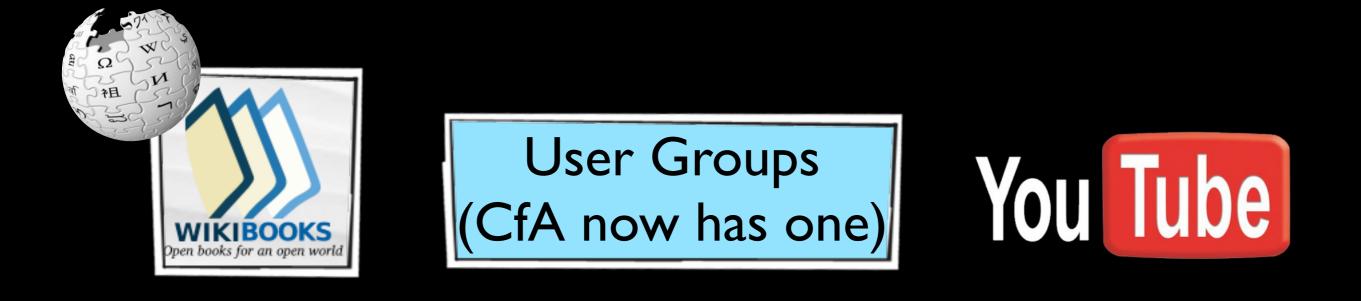
Figure 3 | Schematic illustration of the dendrogram process. Shown is the construction of a dendrogram from a hypothetical one-dimensional

emission profile (black). The dendrogram (blue) can be constructed by 'dropping' a test constant emission level (purple) from above in tiny steps (exaggerated in size here, light lines) until all the local maxima and mergers are found, and connected as shown. The intersection of a test level with the emission is a set of points (for example the light purple dots) in one dimension, a planar curve in two dimensions, and an isosurface in three dimensions. The dendrogram of 3D data shown in Fig. 2c is the direct analogue of the tree shown here, only constructed from 'isosurface' rather than 'point' intersections. It has been sorted and flattened for representation on a flat page, as fully representing dendrograms for 3D data cubes would require four dimensions.



Goodman et al. Nature, 2009

How do we increase the fraction of astronomy researchers who know about these tools?



+Suggestions?!







How do we increase the number of people who create and interlink new tools?

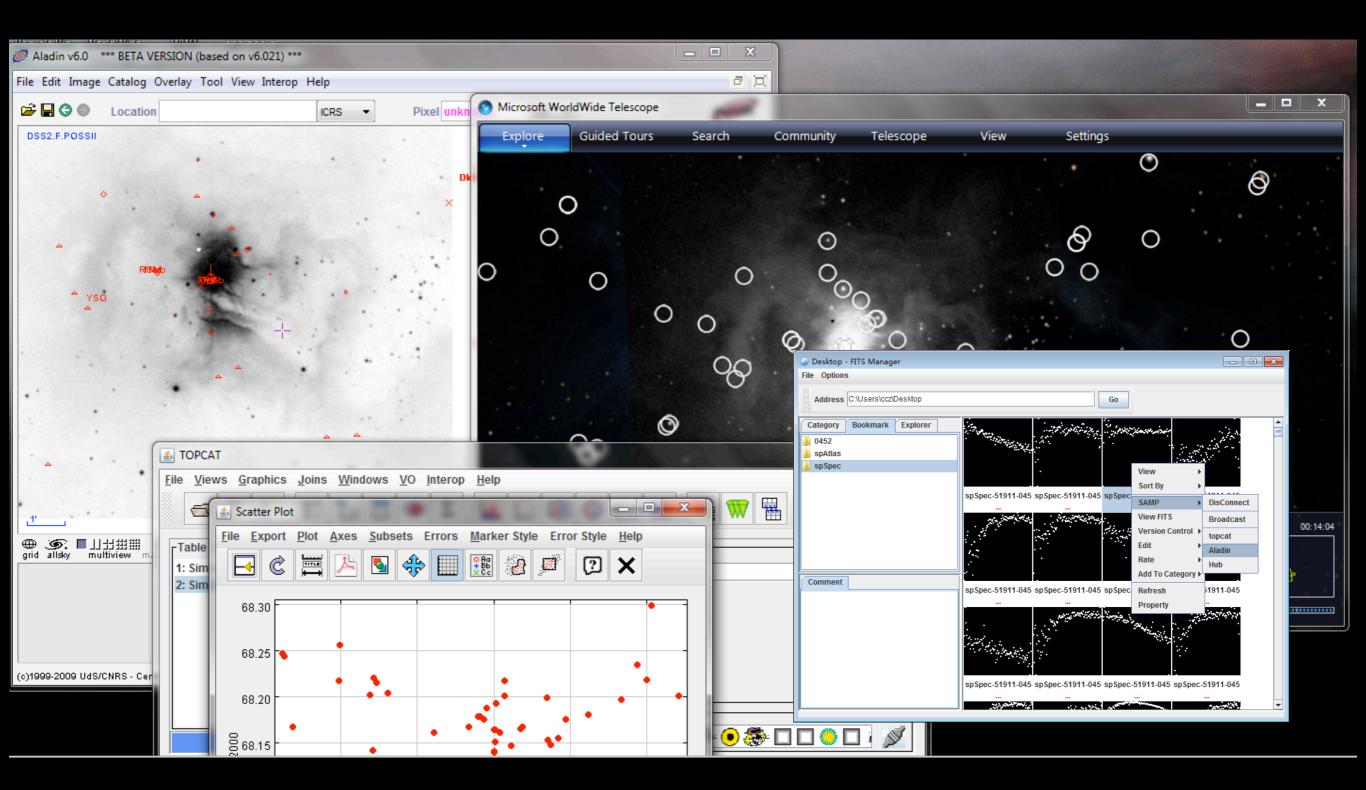


Should implemented through VAO "Associates," WWT Partners, and more.

How do we organize such diverse tools, so as to make them interoperably useful?....

"SAMP" is a great technical start, but offers a very significant user interface challenge. interconnections/modularity

SAMP



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- create intelligent applications which can reason and inference with
- publish resources as Linked Data, externally indexed.
- easily aggregate metrics of interest to publishers, funding agencies
- let others build applications on this substrate using SPARQL querie

We will be carrying out these efforts as part of ADS Labs.

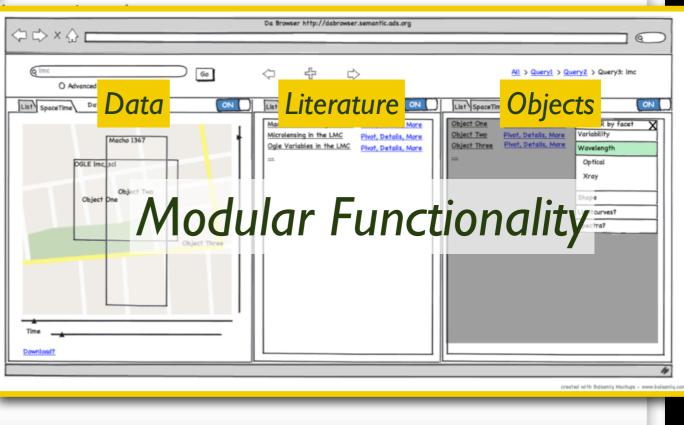
Bootstrapping in ADS Labs

ADS Labs is an effort to put out more forward thinking, somewhat unstat applications will be incubated in ADS Labs before being pushed out to AD

- The results of queries on a bibliographic database will be made available. We will build a user interface on the above (see first image b)
- 2. We will switch to a semantic backend with a SPARQL interface
- 3. Development on Ontologies (which this site details) continues and
- Finally we'll combine the databases so as to have one large semant below.

Examples of Applications

Here are examples of what such applications might look like:



Ongoing "ADS Labs" Work: Rahul Davé, Alberto Accomazzi, Michael Kurtz, AG Thanks to ADS (NASA)/VAO(NASA+NSF)/MSFT funding.

Collaborative Astronomy at University of Washington

Research in a Browser

- "iGoogle" for Astronomy
 - Collections of simple atomic applications (gadgets)
 - Users choose the view they want
 - All gadgets can communicate with each other

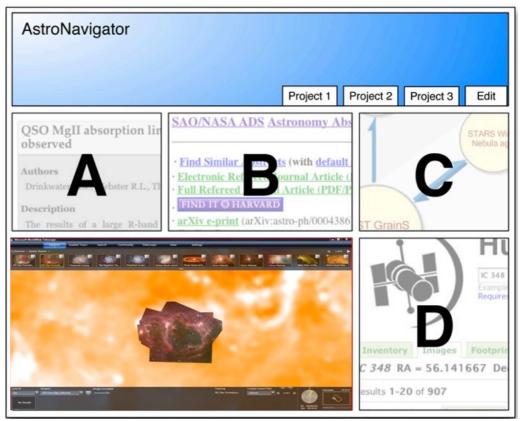
Customizable and sharable

- Users can build and share "mashups"
- Widgets are simple to create
- Widgets call virtual observatory resources

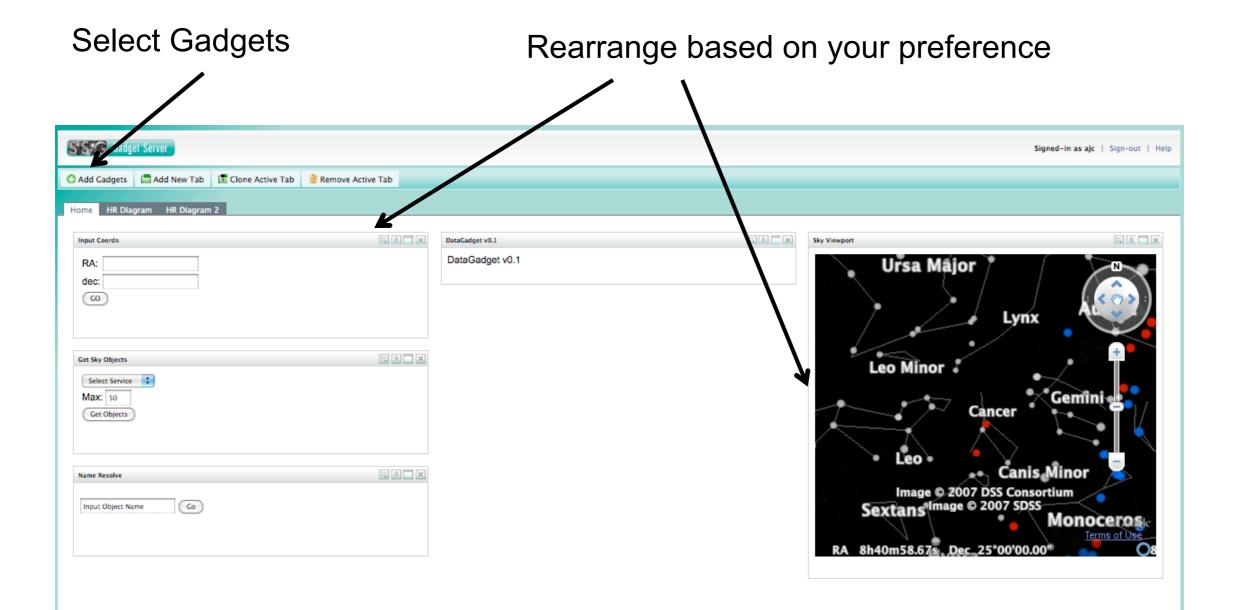
Efficient

- Communication is within the browser (fast)
- Built from javascript (standard)

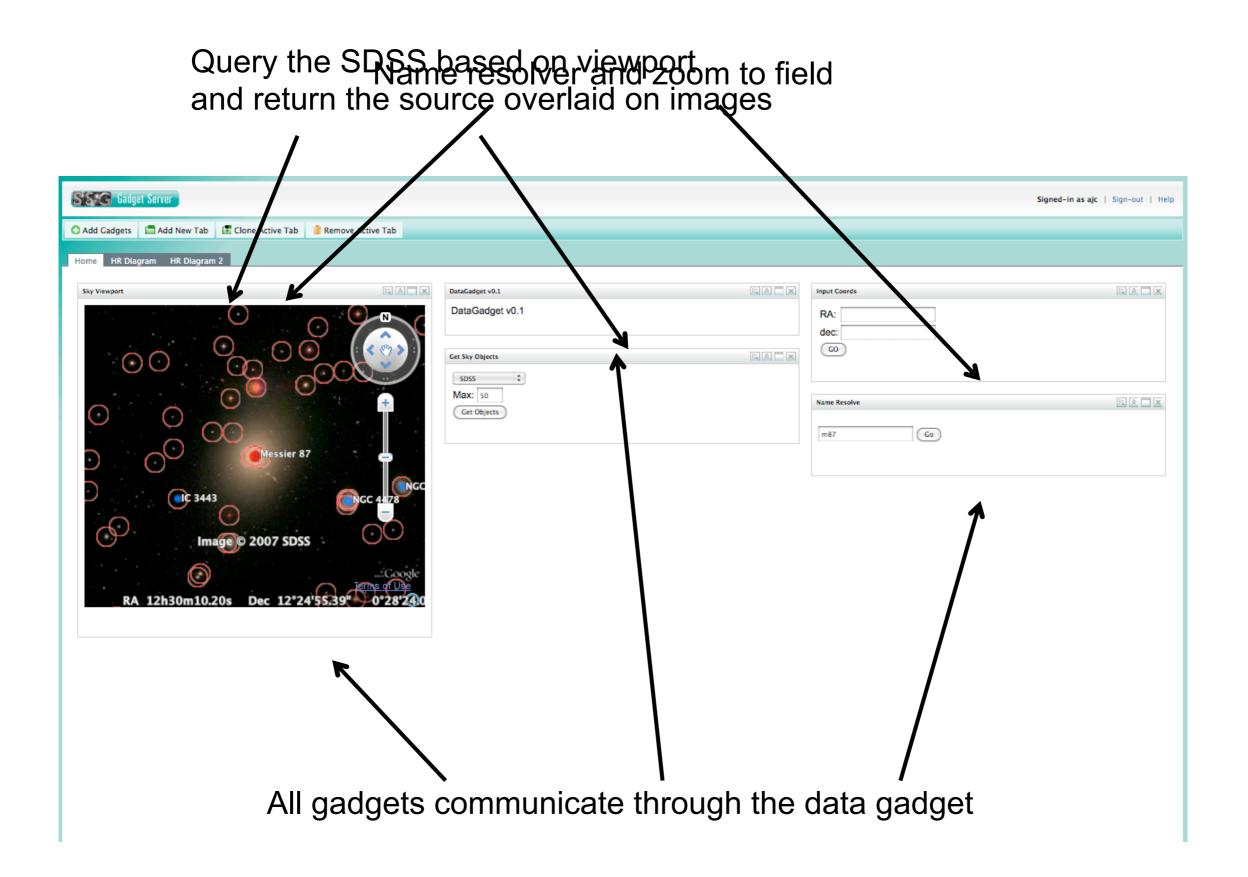
Show Andy Connolly's Movie....







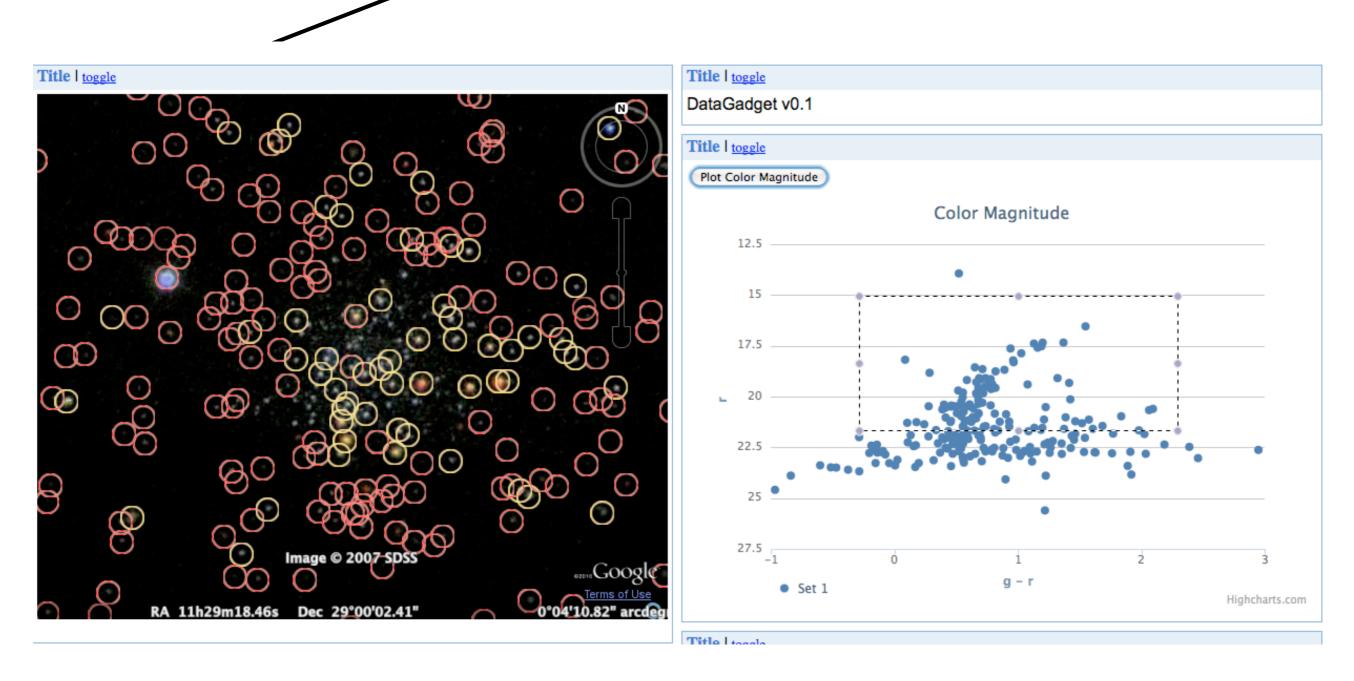






Collaborative Astronomy (Connolly, Gibson, Krughoff, Sayers, Smith 2010)

Create, store and share multiple views of gadgets



Interaction allows selections to be shown on the viewport



Collaborative Astronomy (Connolly, Gibson, Krughoff, Sayers, Smith 2010)



WorldWide Telescope Ambassadors Program

Alyssa Goodman

Harvard University Professor of Astronomy, WGBH Scholar-in-Residence, Microsoft Academic Partner

Annie Valva WGBH Interactive, Director of Research & Development

Pat Udomprasert

WWT Program Coordinator



How?

Using new WWT platform to give experts and learners access to the Universe

WWT Ambassadors Program Recruiting, Vetting, Coordination



data, literature, media

WWT





hosted/ promoted by





"I never knew programs like this could even exist. It's just amazing." –Clarke Middle School 6th grade student

More quotes from Clarke 6th Graders

"Learning about our Universe by actually seeing and exploring it makes it easier to contemplate and more fun."

"You can explore the Universe yourself and you don't always have to only learn from the teacher."

"It gave me a better mental map of the universe."

(And of the 72 surveys we've collected, 71 are positive toward WWT Ambassadors.)